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# Absorption of short-wave radiation in the atmosphere over India in winter

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सार — विश्व मौसम संगठन के विश्व जलवायु कार्यकम के लिए विकिरण अध्ययनों की महता को ध्यान में रखते हुए विभिन्न ऋतुओं में भारत पर सौर और पायिव विकिरणों द्वारा वायमंडलीय स्तंभ के उष्णक/शीतलक का जलवाय संबंधी अध्ययन किया गया । इस शोध पत्न में (भारत में) जनवरी माह में पृथ्वी से सूर्य और विसरित लघु तरग विकिरण द्वारा प्रत्यक्ष लघु-तरंग विकिरण के वायुमंडल में अवचूषण (सोखने) को प्रस्तुत किया गया है । इससे पता चलता है कि आपतित प्रत्यक्ष और विकिरण अर्थात उष्मांक (अधिकम 1360W/mº) का लगभग केवल 25 प्रतिशत भाग वायु मंडल में(345W/mº)अवचृषित होता है और बायु मंडल के ऊपर स्थित कुल मात्रा का मात्र लगभग एक प्रतिप्रत 500 मि. बार तक अवचूषित होता है । 500 और 850 मि. बार के मध्य, इसका लगभग 26 प्रतिशत अवचुषण होता है किन्तु 850 मि. बार और सतह के मध्य यह 34 प्रति-णत होता है। इस प्रकार, वायु मंडल का सबसे नीचे के आधे भाग ( 500 मि. बार के नीचे) बायु मंडल में कुल लघु तरंग विकिरण के अवचृषित भाग का करीब-करीब 60 प्रतिशत सोख लेता है और इसका प्रमुख भाग सबसे नीचे के 1500 मीटर में सोख लिया जाता है ।

ABSTRACT. Considering the importance of radiation studies for the World Climate Programme (WCP)<br>of WMO, a climatic study of warming/cooling of the atmospheric column by solar and terrestrial radiations<br>over India in differ short-wave radiation from the sun and diffuse short-wave radiation from the earth in the month of January (over<br>India) is presented. It shows that only about 25 per cent of the incident direct solar radiation, *viz*., sol

#### 1. Introduction

It is well established that radiation processes play a very effective role in controlling the weather in tropics. For instance, the differential heating of the Indian subcontinent in summer has now been recognised as the main mechanism for triggering the monsoon (Krishnamurti 1979). It is also being recognised that for complete understanding of trade inversions, monsoons, African disturbances and many other weather processes in the tropics, the role of radiation must be fully understood.

In India, climatic study of radiation has been carried out since long based on sunshine records and later on radiation instruments' data (Mani et al. 1963, 1968, 1971, 1981). Recently Mani and Rangarajan (1982) studied the radiation climate of India for the entire year. Their study, however, gives the heat balance at the surface The effect of radiation processes in the atmosphere only. has not been given by them. In fact, studies on radiation budget of the troposphere are very few. Katayama's (1966, 1967, 1972) work which coveres the whole northern hemisphere seems to be one of the few.

In our effort to study the heat balance of the atmosphere over India we present here the climatic study of the short-wave radiation flux only over the country

during winter season. In this study, only the effect of water vapour present in the atmosphere is considered. The contribution of other atmospheric constituents like ozone, CO<sub>2</sub>, aerosols including dust particles which is of significance specially in the lower troposphere is not considered. The study will be extended finally to derive the heat budget of the troposphere over India throughout the year.

### 2. Data

For this study, mean monthly upper air data for the month of January for three years (1978-80) for all the nineteen Indian radiosonde stations (Table 1) published by WMO in the monthly climatic data of the world were used. From these data, three years' mean monthly<br>temperature and dew point values, were worked out for<br>surface, 850, 700, 600, 500, 400, 300, 200 and 100 mb levels for each station using both 0000 and 1200 UT observations. Using these mean temperature and dew point values, mixing ratios at all these levels were picked out with the help of tephigrams for each station.

Dew point temperatures were mostly available upto<br>500 mb level only. In order to estimate dew point temperature values at higher levels, one of the schemes proposed by Katayama (1966) was followed. The dew point curve was extended above 500 mb assuming that it

Computation of absorbed part in incident solar radiation at top of the atmosphere during January (declination for January =22.2°, solar constant  $S_0 \approx 1360$  W/m<sup>2</sup> and time of<br>observation taken as 0060 GMT)



(dew point curve) runs parallel to the air temperature curve upto tropopause and that the mixing ratio remains constant in the stratosphere. The mixing ratio (specific humidity) was thus calculated upto 100 mb level.

#### 3. Numerical calculations

In dealing with tropical weather systems (Krishnamurti 1979) the following calculations are important with respect to position and time:

(i) Rate of warming of the atmosphere by short-wave (solar) radiation,

(ii) Rate of warming/cooling of the atmosphere by long wave (terrestrial) radiation and

(iii) Heat balance of the earth's surface.

In such calculations, prevailing conditions such as the vertical distribution of temperature, moisture, temperature of the underlying surface, nature of the surface, cloud-cover, cloud height, cloud thickness, vertical extension of dust and other atmospheric constituents like ozone, water vapour, CO<sub>2</sub> and nature of the earth's surface are important.

Here we have calculated only the effect of short wave radiation on the atmosphere over India during the month of January, assuming that the atmosphere in this month is mainly cloudless and free of dust in the climatological sense. Effect of surface topography and soil moisture have been ignored. Effects of ozone and other rare constituents of the atmosphere have also not been taken into consideration though important in the lower troposphere. Only the effect of water vapour (moisture content) in the atmosphere has been considered.

It is well known that scattering and absorption are the two main mechanisms which deplete the incident solar radiation in the atmosphere. The solar radiation incident at the top of the atmosphere is represented by solar constant  $(S_0)$  which has the value  $S_0 \approx 1360$  W/m<sup>2</sup>. The expression for scattered and absorbed part of the incident solar radiation is given Katayama (1972) by

Scattered part 
$$
S^s = 0.651 S_0 \times \cos \nu
$$
 (1)

Absorbed part  $S^a = 0.349 S_0 \times \cos \nu$  $(2)$ 

where  $\nu$  is the zenith angle of the sun. In this study 0600 GMT has been taken as time of observations.

## 4. Calculation of incident solar radiation at various levels

Assuming the top of the atmosphere at 0 mb the absorbed parts in the incident solar radiation at the top of the atmosphere  $(S_0)$  were worked out for each station  $(Table 1)$ .

For calculation of the absorbed part of the incident solar radiation through different layers of the atmosphere the atmosphere has been divided into 100 mb thick layers down to 700 mb after which the two layers are 850 mb and surface (average pressure 1000 mb). For these calculations use has been made of concept of optical depth of the atmosphere which is a function of the mixing ratio, pressure and temperature distribution in the atmosphere. It is expressed by the relation (Kuhn  $1963$ :

$$
W_p = \frac{1}{g} \int_{0}^{p} q \left(\frac{p}{p_o}\right)^{0.85} \left(\frac{T_o}{T}\right)^{0.5} dp \tag{3}
$$

where  $W_p$  is the 'optical depth' of the atmosphere at a reference level  $(p)$  estimated from the top of the atmosphere  $(p\neq 0)$ .  $p_0$ ,  $T_0$  are the normal pressure and absolute temperature (1013.2 mb and 288°K) at the surface, and  $p$ ,  $T$  are the pressure and temperature of the reference layer.  $q$  is the mixing ratio of the layer and  $dp$  is the depth of the layer in millibars. With the help of this equation, optical path length at the pressure levels, 100, 200, 300, 400, 500, 600, 700, 850 and the surface were calculated.

For calculation of absorption in incident solar radiation in each layer, use has been made of Joseph's (1966) absorptivity function  $[(A(W)]$  which gives the value of incident radiation depleted by the absorbing constituent, viz., water vapour. Here :

$$
A(W)_p = 0.271 \ (W_p \ \sec \ \nu)^{0.303} \tag{4}
$$

where " $W_p$  sec  $v$ " is the effective path length through which solar radiation has passed.

Hence, the incident radiation reaching a reference level  $p$  is written as :

$$
(DSR)_p = S^a \left[ 1 - A \left( W \right)_p \left( W_p \text{ sec } v \right) \right] \tag{5}
$$

where  $S<sup>a</sup>$  is the absorbed part in the solar radiation incident at the top of the atmosphere (0 mb level) and " $W_p$  sec v" the effective path length in the layer. Using<br>this expression, the amounts of absorbed part (S<sup>a</sup>) of direct solar radiation (DSR) reaching various levels of

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#### **TABLE 2**

The amout of absorbed part  $(S^q)$  in solar radiation reaching at various levels over India during the month of January ( $W/m<sup>2</sup>$ ) from top of the atmosphere



the atmosphere from the immediate upper levels were worked out for the whole atmosphere from top to surface and level by level for the month of January as shown in Table 2. This table shows that the absorbed part of solar radiation gets very much absorbed by water vapour on reaching 850 mb level with respect to the top of atmosphere.

# 5. Calculations of diffuse solar radiation at various levels

On reaching the earth's surface a part the incident solar radiation gets reflected back into the atmosphere as reflected radiation depending on the albedo of the earth surface. In this case we have assumed the earth surface to be covered with green grass for which average surface albedo  $*_a$  is taken to be 0.2. Thus the incident solar radiation reflected back to space by the earth's surface will be given by :

$$
(DFR)_{\text{surface}} = S_0^a [1 - A W_0^* (W_0 \sec \nu)]^* a_s \quad (6)
$$

The reflected short wave radiation experiences in general a longer path length than the direct solar radiation. Joseph (1966) suggested that the path length of the reflected radiation be increased by a factor of 1.66. Then the amount of reflected short wave radiation reaching a reference pressure level  $(p)$  from the earth's surface may be given by :

$$
(\text{DFR})_p = S_0^a \left[ 1 - A W_0 (W_0 \sec \nu)^* \alpha_s \right] \times \left[ 1 - A W_p \left\{ 1.66 (W_0 - W_p) \right\} \right] \tag{7}
$$

In this expression the first part accounts for the total reflected solar radiation for the earth's surface (albedo) and the second part represents the absorption of reflected radiation due to the increased path length between the earth's surface and the reference level  $(p)$ .

With the help of this expression, reflected solar radiation energy from surface reaching at  $850$ ,  $700$ ,  $600$ ,  $500$ ,  $400$ ,  $300$ ,  $200$  and  $100$  mb levels has been calculated for each station. In this case reflected radiation at the lower level has been considered as the source for the higher level. The result of this computation has been given in Table 3.

## 6. Calculation of net downward short-wave radiation

The net downward solar radiation flux (SRF) at any level  $(p)$  in the atmosphere was obtained as the algebraic difference between the direct solar radiation and the reflected radiation at that level :

$$
(SRF)_p = (DSR)_p - (DFR)_p \tag{8}
$$

The net downward flux of solar radiation calculated for 19 stations over the country is given in Table 4.

# 7. Analysis

## 7.1. Direct solar radiation

The analysis of the data shows that only about 25 per cent of the incident solar radiation at the top of the atmosphere gets absorbed in the atmosphere, i.e., only 345 W/m<sup>2</sup> out of the incident solar energy of about 1360 W/m<sup>2</sup> (solar constant) is absorbed in the atmosphere.

The absorbed part of energy at the top of the atmosphere is almost undiminished upto 500 mb level  $(99\%)$ below which attenuation starts. Still nearly  $73\%$  of the energy reaches 850 mb level. The real attenuation starts in the lowest part of the atmosphere, i.e., below 850 mb due to large amount of moisture present in the lowest layers. The atmosphere above 500 mb is almost transparent to incident solar radiation with absorption just

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#### TABLE 3

Amount of reflected solar radiation reaching various levels (W/m<sup>2</sup>) from earth's surface after reflection during the month of January over India





Fig. 1. Direct short-wave radiation (DSR) and net amount of downward short-wave radiation flux (SRF) at various levels over India during the month of January

about  $1\%$  of the energy received at the top of the atmosphere. Below 500 mb the absorption is higher upto  $850 \text{ mb} (26\%)$  but between  $850 \text{ mb}$  and the surface<br>it is maximum. Fig. 1 brings out very clearly how fast the attenuation takes place in the first 1500 m of the atmosphere.

### 7.2. Net solar radiation

Table 4 gives the net solar radiation reaching various levels of the atmosphere after taking into account the loss due to reflection. It is also shown in Fig. 1. From a comparison of direct and net solar radiation curves (Fig. 1), it is apparent that loss due to reflection is high in lower layers and very little or even slight increase in higher layers which may be due to absorption by dust and aerosol particles. In fact, it is quite high upto 850 mb (Table 3) level, then decreases slowly upto 700 mb level and becomes almost *nil* or shows<br>slight rise above 600 mb. Depletion is again due to absorption of a part of reflected short-wave radiation from the surface by the moisture in lower levels.

## 7.3. Spatial distribution of net solar radiation over India

Fig. 2 shows the spatial distribution of net solar radiation at surface, 850, 700, 500 and 100 mb levels.<br>It can be seen that at 100 mb the net incident solar radiation is almost zonal decreasing from south to north as is natural with sun being south of equator

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Figs. 2(a-d). Net amount of downward flux of short-wave radiation (W/m<sup>a</sup>) during January

in this month. This condition continues right upto 500 mb with little change. Below 500 mb, the attenuation due to moisture becomes apparent, first at southern coastal stations, particularly at the west coast. At the surface levels, however, the picture is quite interesting. The landmass of central and north India receives the maximum radiation due to cloudless condition and coastal belts the least. The decrease in coastal belts is due to relatively high amount of moisture at coastal stations in lower levels when inland stations are almost dry.

An analysis of inland and coastal stations separately shows that the average net solar radiation reaching surface for land stations (Hyderabad, Nagpur, Ahmedabad, Jodhpur, Delhi, Srinagar, Lucknow and Gauhati) is 130W/m<sup>2</sup> and for coastal stations (remaining stations in Table 2) it is 81W/m<sup>2</sup>.

#### 8. Conclusion

The following broad conclusions can be derived from this study :

(i) Only about 25% of the incident direct solar radiation, (i.e., solar constant of about 1360 W/m<sup>2</sup>), is absorbed by the water vapour of atmosphere  $(345 \text{ W/m}^2)$ .

(ii) Atmosphere is almost transparent to incoming solar radiation upto 500 mb levels.

(iii) Between 500 mb and 850 mb levels the absorption is  $26\%$  and between 850 mb and surface, it is  $34\%$ .<br>Thus,  $60\%$  of the solar radiation at the top of the atmosphere gets absorbed in the lowest half of the atmosphere below 500 mb, of which about 58% in the lowest 1500 metres.

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#### TABLE 4

Net amount of downward flux of solar radiation ( $W/m^2$ ) at various levels of the atmosphere over India during the month of January



(iv) Approximately  $40\%$  (i.e., 135 W/m<sup>2</sup>) of the direct solar radiation entering the atmosphere (345 W/m<sup>2</sup>) ultimately reaches the surface on an average. It is just about  $10\%$  of the solar constant.

 $(v)$  The reflected radiation (reflected solar radiation from the surface) decreases upward layerwise. The loss is high in lower levels.

(vi) The net downward solar radiation flux decreases from south to north over India in January upto 850 mb levels. At the surface, however, the picture becomes quite distorted with coastal stations showing much less incident short-wave radiation than the inland stations. The average incident radiation at coastal stations is 81 W/m<sup>2</sup> whereas over inland stations it is 130 W/m<sup>2</sup>. This is because of strong absorption by water vapour present in lower levels at coastal stations, inland stations being drier at the surface in this month.

(vii) The water vapour emerges as a strong absorber of incident solar radiation both direct as well as reflected.

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